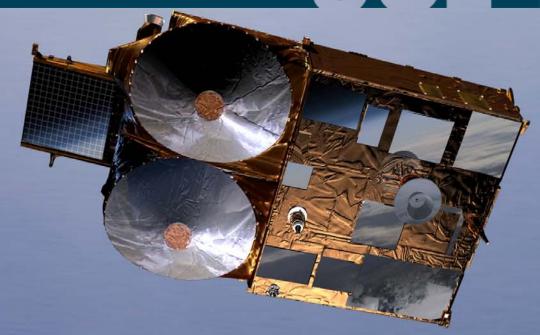
LUCL

Precise Estimates of Ocean Surface Parameters from CryoSat



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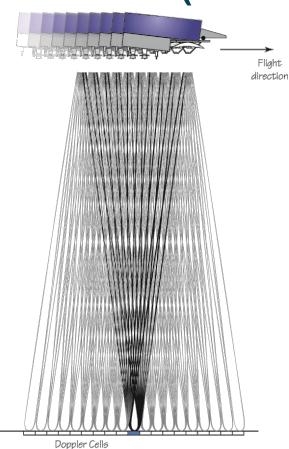
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- 2. European Space Agency
- 3. NOAA

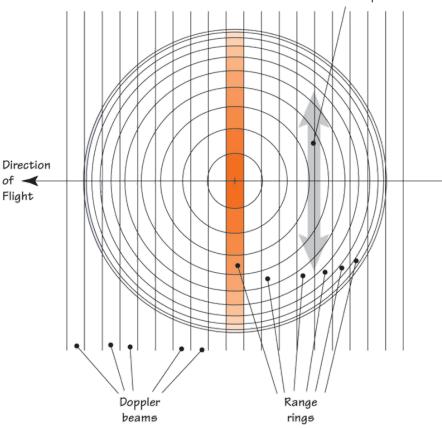


interferometer phase

Mission concept: Low Resolution Mode (LRM), Synthetic Aperture Radar (SAR) mode and SAR Interferometric (SARIn) mode.

Angles inferred from





[Wingham et al., Adv. Sp. Res., 2006]







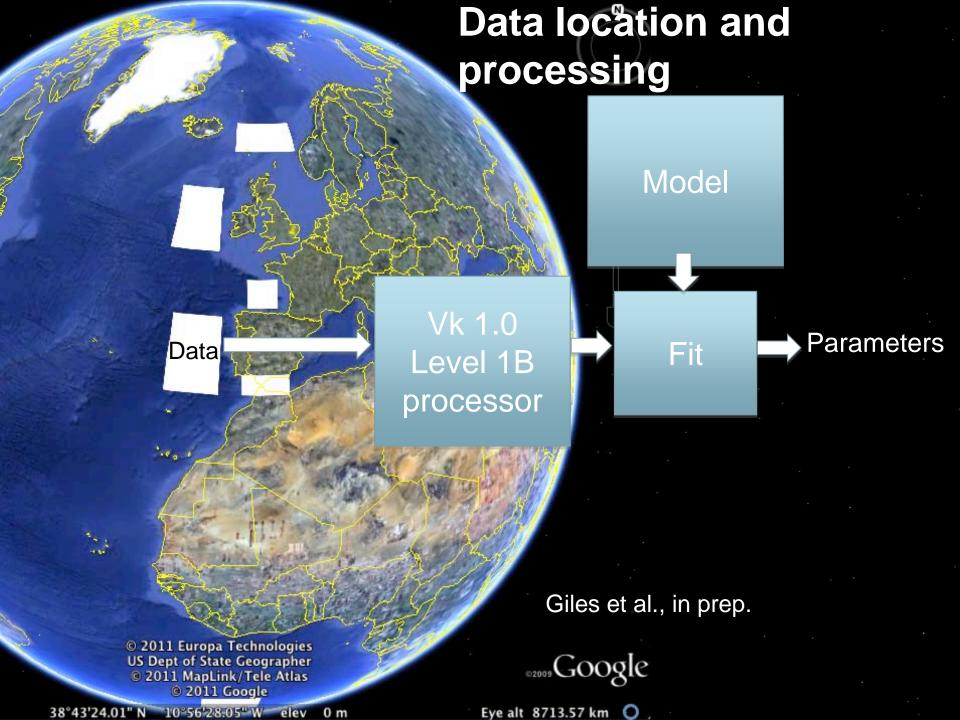
CryoSat-2 SAR mode over the open ocean

 Can we see an improvement in the precision of the range measurement and significant wave height in the CryoSat-2 data?

We note here that CryoSat-2's SAR mode is not designed to give the maximum number of looks possible in SAR mode, this experiment is designed to directly compare measurements taken in SAR mode and pulse limited mode. Raney (2012) has described the instrument set up required to obtain the maximum number of looks.









Model

$$p_r(\tau) = p_t(\tau) * p_z(\tau) * I(\tau)$$
 General expression for the mean each power (Brown, 1977)

$$p_t = p_0 g_p \mathrm{sinc}^2(\pi B \tau)$$
 Transmitted power pulse from CryoSat-2 $p_z = \frac{1}{\sqrt{2\pi\sigma}} e^{-\frac{1}{2}(\tau/\sigma)^2}$ Surface elevation pdf

I depends on the backscatter (σ^0), antenna gain and range from the radar to the surface

One-way gain of a synthetic beam

Integration around a range ring
$$\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{d\theta} \int_{0}^{2\pi} \frac{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{\int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2} \int_{0}^{2\pi} d\theta \frac{d(\rho_{k} \cos \theta - \xi_{k})}{2}$$

$$\cdot \exp \left[-2 \left(\frac{(\rho_k \cos \theta - \mu)^2}{\gamma_1^2} + \frac{(\rho_k \sin \theta - \chi)^2}{\gamma_2^2} \right) \right]$$
pitch Elliptical antenna pattern roll

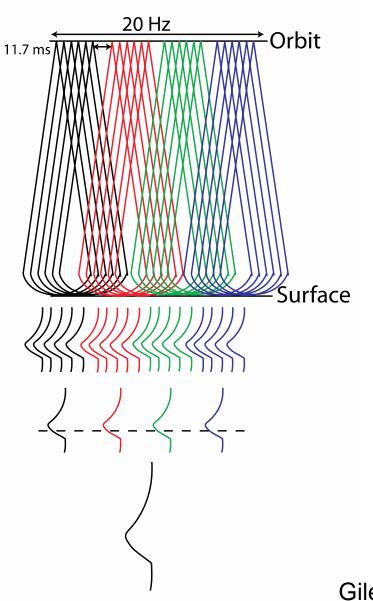
Wingham et al., 2004

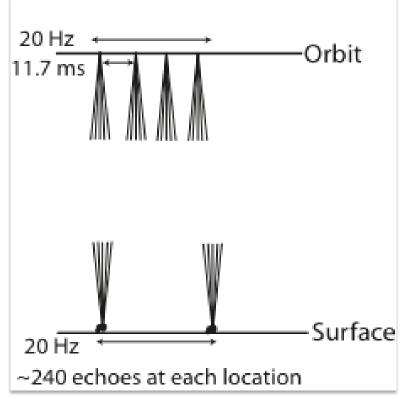


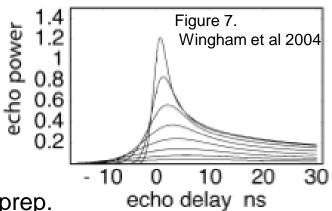


Pulse limited data

SAR mode data • UCL



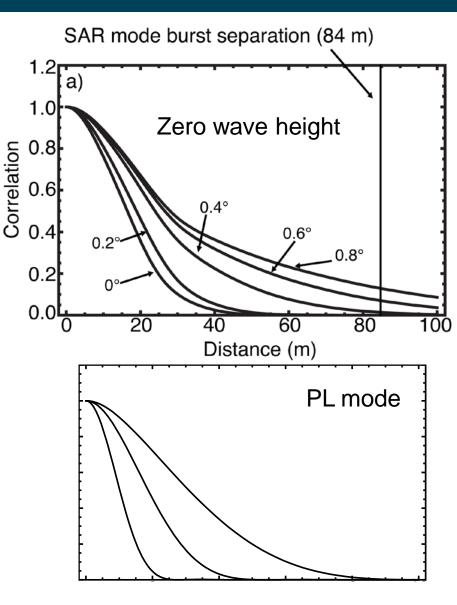


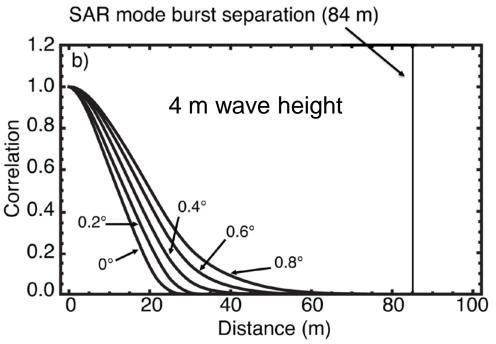


Giles et al., in prep.

Power correlation coefficient







SAR mode correlations as a function of the look angle

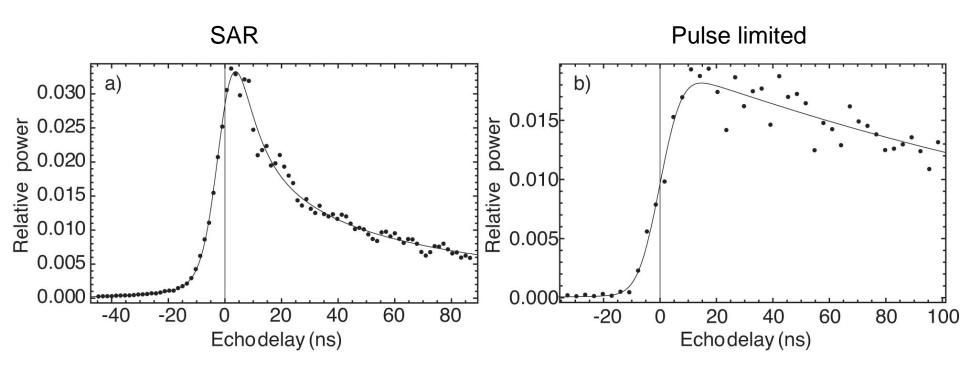
PL correlations and a function of wave height 0, 2 and 4 m wave heights. The separation of the echoes is 0.41 m







Examples of the model fit to the data



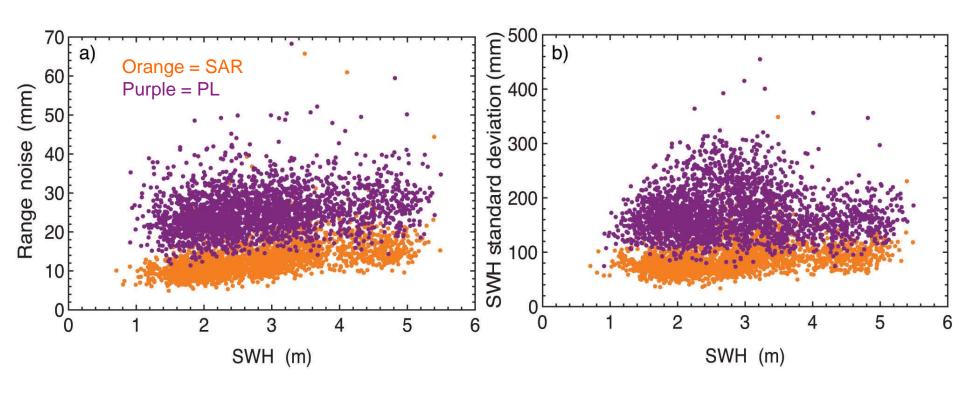
Giles et al., in prep.







Results: Precision in the range and SWH



- SAR range noise at 1 Hz and SWH ~2 m is 1.1 cm
- PL range noise at 1 Hz and SWH ~2 m is 2.3 cm
- N.B. CryoSat LRM range noise is 1.3 cm [Smith and Scharoo, OSTST 2011]
- SAR SWH noise at 1 Hz and SWH ~2 m is **7.5 cm**
- PL SWH noise at 1 Hz and SWH ~2 m is 17.0 cm

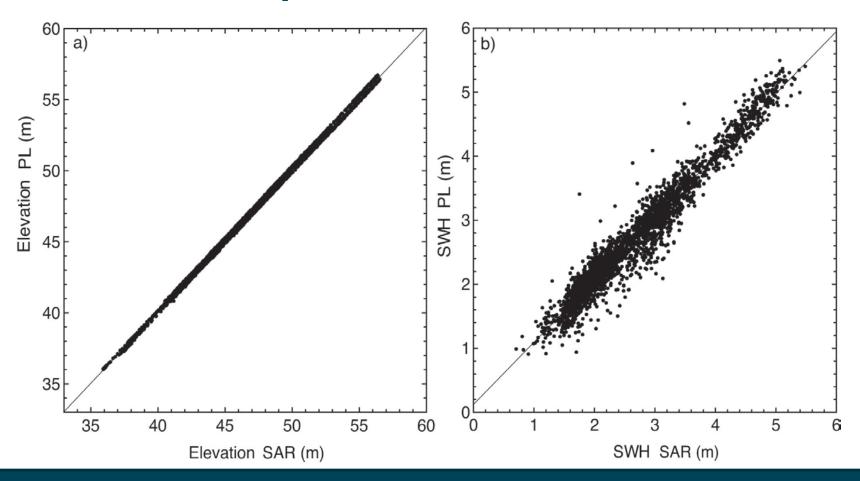
Giles et al., in prep.







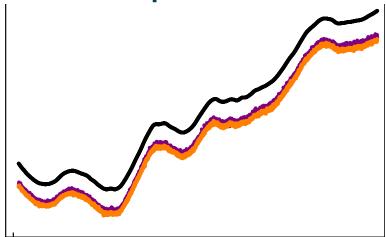
Results: Comparison between modes



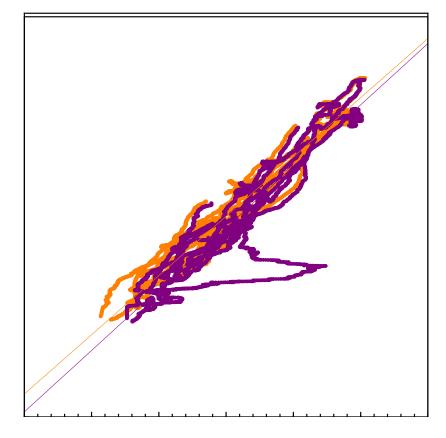
Regression of the a) elevation and b) SWH from the SAR and PL modes. The fit to the data in a) is given by Elevation_{PL}= 1.002Elevation_{SAR}+0.009 and in b) by SWH_{PL}=0.931SWH_{SAR}+0.349. Including the pitch and roll brought the values closer together



Results: Comparison with the EGM08 geoid and SWH from ECMWF

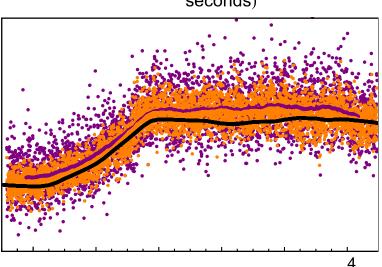


ECMWF_SWH=0.407244 + 0.872486 SAR_SWH ECMWF_SWH=0.144844 + 0.903942 PL_SWH



Orange) PL (purple) (m)

seconds)



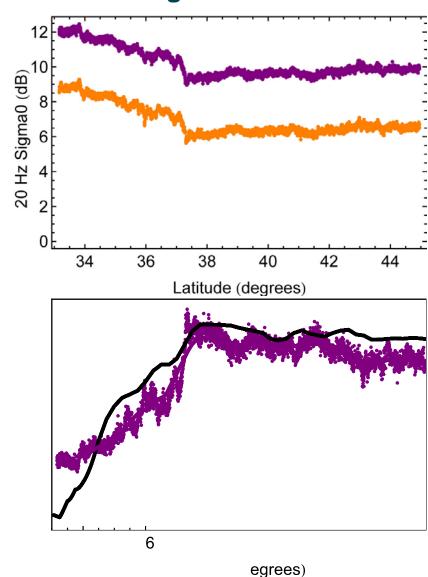
SWH data from Dr. Saleh Abdalla @ ECMWF



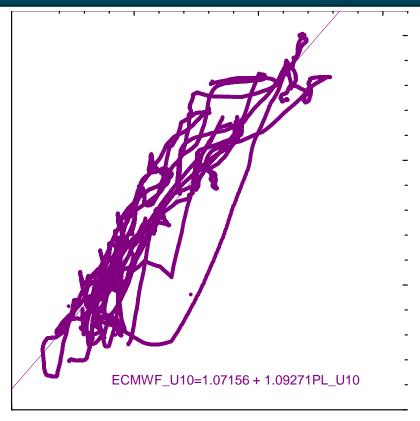




Results: Sigma-zero and wind speed from ECMWF



Sigma-zero was converted to wind speed using Abdalla's formula for Envisat [Abdalla (2007), Proc. 'Envisat Symposium 2007']



PL U10 windspeed (m)

Wind speed data from Dr. Saleh Abdalla @ ECMWF







Conclusions

- We have directly compared the precision in the measurement of oceanographic parameters by re-tracking the data using a semi-analytical model.
- We have demonstrated that SAR processing can provide a two fold improvement to the precision compared to PL processing.
- The SAR mode precision of 1.1 cm is similar to the LRM precision of 1.3 cm found by Smith and Scharroo (OSTST 2011).
- The oceanographic parameters from CryoSat-2 compare well to the EGM08 geoid and data from ECMWF.
- There is a fixed offset in the sigma-zero estimates from our SAR data that may originate in the UCL SAR processing.
- For future missions, an approximate three fold improvement could be made on the precision achieved by CryoSat's SAR mode and is described by Raney (2012).



